

Journal of
Astronomical Telescopes,
Instruments, and Systems

AstronomicalTelescopes.SPIEDigitalLibrary.org

Flux-pinning mechanisms for improving cryogenic segmented mirror performance

Jessica Gersh-Range
William R. Arnold, Sr.
David Lehner
H. Philip Stahl

Flux-pinning mechanisms for improving cryogenic segmented mirror performance

Jessica Gersh-Range,^{a,*} William R. Arnold Sr.,^b David Lehner,^{c,†} and H. Philip Stahl^c

^aCornell University, 127 Upson Hall, Ithaca, New York 14853, United States

^bDefense Acquisition, Inc., Jacobs ESSSA Group, Huntsville, Alabama 35806, United States

^cNASA Marshall Space Flight Center, Huntsville, Alabama 35812, United States

Abstract. Although large cryogenic space telescopes may provide a means of answering compelling astrophysics questions, the required increase in the primary mirror diameter presents technical challenges. Larger primaries are more flexible, and cryogenic mirrors are typically very lightly damped—the material damping is negligible, and common damping methods break down. To address these challenges, we propose placing flux-pinning mechanisms along the edges of adjacent mirror segments. These mechanisms consist of a collection of magnets and superconductors, and like flexures, they preferentially allow motion in specific degrees of freedom. Motion in nonpreferred degrees of freedom is resisted by a force analogous to a damped spring force, and the stiffness and damping can be adjusted independently. As an example, we consider simple mechanisms consisting of an inexpensive magnet and a single superconductor. These mechanisms provide increasing resistance as the magnet and superconductor—or mirror segments attached to each—come closer to colliding. These mechanisms, with typical stiffness and damping values on the order of 5000 N/m and 5 kg/s, respectively, also provide modest improvements to the mirror performance. Greater gains can be achieved by using stronger magnets or smaller separations, or by placing nonmagnetic conductive materials near the mechanism.

© The Authors. Published by SPIE under a Creative Commons Attribution 3.0 Unported License. Distribution or reproduction of this work in whole or in part requires full attribution of the original publication, including its DOI. [DOI: [10.1117/1.JATIS.1.1.014001](https://doi.org/10.1117/1.JATIS.1.1.014001)]

Keywords: segmented mirror; space telescope; flux pinning; cryogenic; edgewise connectivity; stiffness; damping; flexure.

Paper 14003P received Mar. 3, 2014; revised manuscript received Jun. 10, 2014; accepted for publication Jul. 9, 2014; published online Oct. 24, 2014.

1 Introduction

Large cryogenic space telescopes may provide a means of answering several compelling astrophysics questions, but the required increase in the primary mirror diameter presents numerous technical challenges. Proposed investigations of early star formation, planetary system evolution, and the presence of large organic molecules in interstellar disks, for example, require a 10- to 16-m class far-IR to submillimeter space telescope; in this wavelength range, atmospheric extinction precludes ground-based measurements, and existing space telescopes suffer from source confusion.^{1–4} It is also expected that high-resolution far-IR observations will lead to new discoveries, possibly concerning the emergence of cosmic structure.^{4,5} However, for an observatory to obtain an angular resolution in the far-IR that is comparable to Hubble's resolution in the visible, its diameter must be on the order of 1 km.⁵ By comparison, Herschel, the largest space telescope to operate in the far-IR to submillimeter range, has a primary diameter of 3.5 m.⁶ As the primary mirror diameter of future cryogenic observatories increases, maintaining a stable wavefront becomes increasingly challenging since the first natural frequency decreases as the diameter squared, the material damping is negligible, and other common damping methods break down at low temperatures.^{7–11}

Although large stiff precision structures are considered an enabling technology for large cryogenic mirrors,¹² an alternative approach to increasing mirror stability is to use an edgewise-connected architecture. In this approach, mechanisms analogous to damped springs are placed along the edges of the primary mirror segments. The stiffness and damping contributions from the mechanisms reduce the requirements for the supporting structure, and if the mechanisms are sufficiently stiff, the segmented mirror performs comparably to a monolith even if the mechanisms are the only structural connections to the segments.¹³ While the mechanisms can be a collection of damped springs or any other device with similar behavior, flux-pinning mechanisms are uniquely suited for cryogenic mirrors.

Unlike mechanical devices, which can have problems with lubrication, coefficient of thermal expansion (CTE) matching, and thermal snap, flux-pinning mechanisms operate best at cryogenic temperatures. These passively stable, noncontacting mechanisms consist of a collection of magnets and type II superconductors and require only low temperatures; no power is needed other than the minimal amount, if any, necessary for cooling. Like a flexure, a flux-pinning mechanism preferentially allows motion in specific degrees of freedom, which depend on the mechanism design, as described in Sec. 2. Motion in the nonpreferred degrees of freedom is resisted by a force analogous to a damped spring force, and the stiffness and damping can be adjusted independently. These mechanism properties depend on the choice of magnets, the separation between the magnets and superconductors, and the presence of nonmagnetic conductive materials, such as aluminum.

*Address all correspondence to: Jessica Gersh-Range, E-mail: jag389@cornell.edu

†Retired.

As an example, we consider simple mechanisms consisting of an inexpensive magnet and a single superconductor separated by distances on the order of 1 mm (Secs. 3 and 4). These mechanisms can be trained to follow a particular displacement pattern when loaded, and they provide increasing resistance the closer the magnet and superconductor—or mirror segments attached to each—come to colliding. Typical stiffness and damping values are on the order of 5000 N/m and 5 kg/s, respectively (Sec. 3). As shown in Sec. 4, mechanisms with these values provide modest improvements to the performance of an edgewise-connected mirror. Greater gains can be achieved by using stronger magnets or smaller separations, or by placing nonmagnetic conductive materials near the mechanism.

2 Flux-Pinning Mechanisms

Flux pinning, a physical interaction between a magnet and a type II superconductor, is analogous to a noncontacting damped spring force. This force is due to the presence of impurities that allow the magnetic field to penetrate into the superconductor material. As the superconductor cools below its critical temperature [approximately 90 K for yttrium barium copper oxide (YBCO)¹⁴], the magnetic field lines are “frozen” in place; motions that change the magnetic field distribution inside the superconductor induce supercurrents that oppose the change. As a result, the relative orientation and position of the magnet and superconductor are passively stabilized in every direction that has a magnetic field gradient.

A useful conceptual model is the frozen-image model, which explains the forces between a magnet and a type II superconductor by replacing the superconductor with two image magnets.¹⁵ The first image magnet, called the mobile image, is a reflection of the permanent magnet across the superconductor surface (Fig. 1). This image moves as the permanent magnet moves, and since the two magnets have opposite moments, the interaction between them is repulsive. The second image

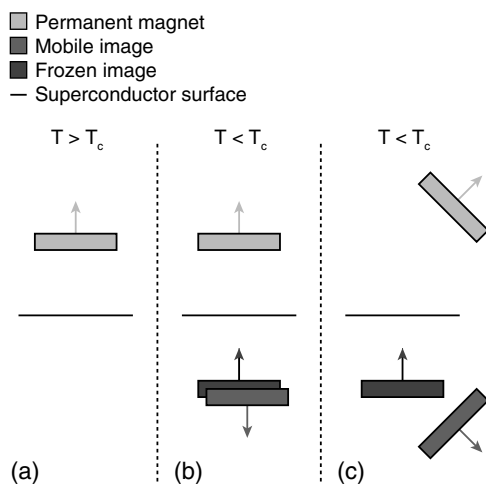


Fig. 1 Image model for flux pinning. In this model, flux pinning is described as the interaction between the permanent magnet and two image magnets that form as the superconductor cools below its critical temperature (a, b). The mobile image is a mirror image of the permanent magnet, and it moves as the permanent magnet moves (c). The frozen image is almost a mirror image of the permanent magnet at the moment the superconductor cools below its critical temperature; the orientation and position are mirror images, but the dipole vector is the same. The frozen image does not move. The force on the permanent magnet is the sum of the forces due to the image magnets.

magnet, called the frozen image, is stationary. This image is almost a perfect reflection of the permanent magnet at the time the superconductor cooled below its critical temperature; the orientation and position of the frozen image are reflections of the orientation and position of the permanent magnet at this time, but the magnetic moment is in the same direction as that of the permanent magnet. The force on the permanent magnet is the sum of the forces due to each image magnet. As a result, the permanent magnet is in equilibrium when it is in its initial orientation and position since the forces from the images cancel. As the magnet moves closer to the superconductor, the force from the mobile image dominates, and the magnet is repelled. Similarly, as the magnet moves further from the superconductor, the force from the frozen image dominates, and the magnet is attracted.

For small motions, the force between the magnet and superconductor is analogous to a damped spring force, with the stiffness and damping determined by a variety of factors including the choice of magnet and superconductor, the separation during cooling, and the presence of conductive materials.^{16–18} For a cylindrical magnet levitated over a cylindrical superconductor, the stiffness for motions perpendicular to the magnet-superconductor interface has been determined empirically to be approximately twice the stiffness for motions parallel to the interface,^{19,20} and the stiffness increases nearly exponentially as the cooling separation decreases.¹⁶ Stiffer interfaces result from using stronger magnets, stronger superconductors, and smaller separations. Additionally, the amount of damping can be increased independently by placing nonmagnetic conductive materials, such as aluminum, near the interface. The inherent damping arises from hysteretic losses,^{21,22} and placing nonmagnetic conductive material nearby increases the total damping by providing a source of eddy current damping. Previous experiments have demonstrated that placing aluminum near the magnet and superconductor can increase the damping substantially, altering the response from imperceptibly damped to strongly overdamped.¹⁶

A flux-pinning mechanism is a configuration of magnets and superconductors that, like a flexure, preferentially allows motion in specific degrees of freedom. Since the mechanism stiffness depends on the magnet-superconductor cooling separation, these mechanisms fall into two broad categories: low-stiffness mechanisms that prioritize larger separations, and low-separation mechanisms that prioritize higher stiffness. Prior research into flux-pinning mechanisms has concentrated on designing mechanisms that serve as joints between spacecraft modules.^{16,23,24} While these mechanisms belong in the first category since large separations are desired, similar techniques can be applied to designing optomechanical interfaces, which typically require high stiffnesses in order to control motion to a fraction of a wavelength.^{25–28}

Since flux pinning resists changes to the magnetic field distribution inside the superconductor, one design approach is to shape the magnetic field so that it is constant in directions where motion is desirable. For example, to allow only translation along a line, like a parallel-blade flexure, one mechanism design consists of a cubical superconductor pinned between two long cylindrical magnets [Fig. 2(a)].[‡] Since a line between and parallel to the magnets is the only direction without a magnetic field gradient, it is the only direction in which the superconductor can move freely; the geometry of the superconductor restricts

[‡]This configuration corresponds to the prismatic joint described in Refs. 16 and 24.

rotation about this line. Similarly, to allow only rotation about an axis, like a cross-blade flexure, the flux-pinning mechanism consists of a small cylindrical magnet pinned above a superconducting disk, as shown in Fig. 2(b). Since the line connecting the magnet and superconductor is an axis of symmetry for the magnetic field, rotations about this axis are unconstrained. Motions in any other direction, however, are opposed by a restoring force since there is a magnetic field gradient.

An alternative approach to flux-pinning mechanism design is to combine field shaping with mechanical constraints. As an example, a flux-pinning mechanism analogous to a wire flexure consists of a cylindrical superconductor pinned to a spherical magnet mounted inside an inverted cone [Fig. 2(c)]. This mechanism allows motion in five degrees of freedom. As before, the line connecting the magnet and superconductor is an axis of symmetry for the magnetic field, so the superconductor is free to rotate about this axis without disturbing the magnet. However, a springlike restoring force opposes any motions that change the relative orientation and position of the magnet, such as motion along the line connecting the two. As a result, rotating the superconductor about the cone will cause the magnet to rotate correspondingly since it is free to move within the cone; radial translation is the only motion the mechanism resists.

For the edgewise-connected mirror analysis that follows, we have selected the mechanism described in Fig. 2(b). This mechanism consists of a 56-mm single-domain melt-textured YBCO disk¹⁴ paired with a neodymium disk magnet, and the

cooling separation varies from 0.5 to 2 mm. To represent this mechanism as a collection of collocated damped springs, it is assumed that all of the important dynamics can be captured by considering only the relative translations between the magnet and superconductor. While this mechanism does resist bending to some extent, it is assumed that the resistance from a single mechanism is negligible. Instead, bending stiffness is added by using pairs of mechanisms, as shown in Fig. 3, separated by a distance d_r .

In this model, there are two rows of mechanisms along each segment edge, and each pair of mechanisms is represented by a set of four collocated single-degree-of-freedom damped springs. One of these damped springs corresponds to translation perpendicular to the superconductor surface, with stiffness k_{\perp} , and two of the springs correspond to translation parallel to the superconductor surface, with stiffness k_{\parallel} . The remaining spring corresponds to bending, with stiffness $k_b = (1/2)k_{\perp}d_r$. Since flux pinning is approximately twice as stiff for motion

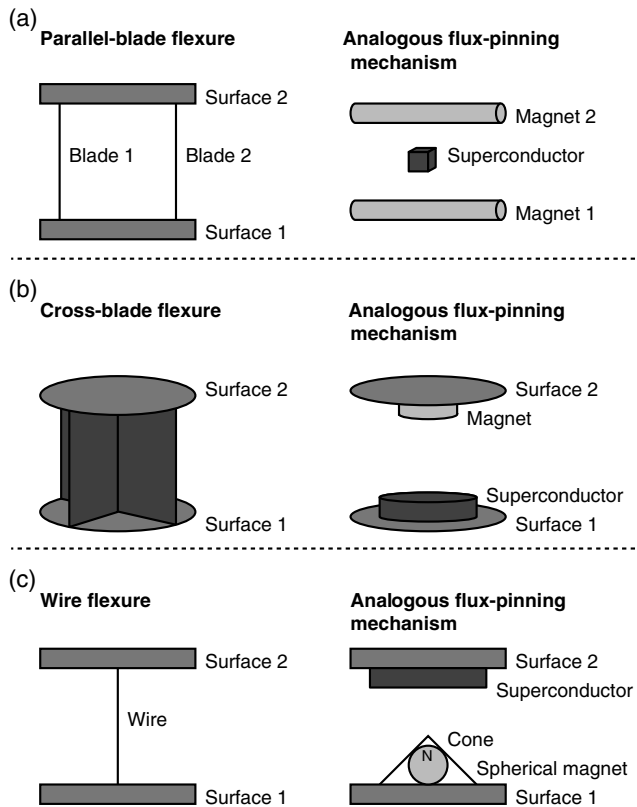


Fig. 2 Flux-pinning mechanisms and the corresponding flexures. Like a flexure, a flux-pinning mechanism preferentially allows motion in specific degrees of freedom. These degrees of freedom depend on the mechanism design, and there are two basic design approaches: shaping the magnetic field so that there is no change in the degrees of freedom in which motion is desired (a, b), and combining field shaping with mechanical constraints (c).

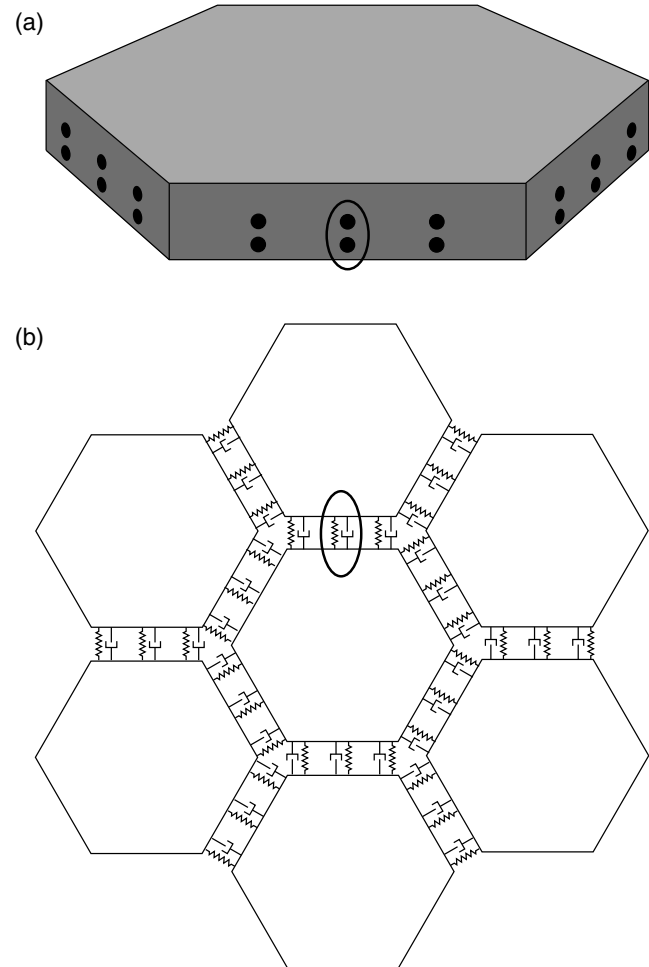


Fig. 3 The mechanism model. As an example, we consider the case of an edgewise-connected mirror using the flux-pinning mechanisms shown in Fig. 2(b). While these mechanisms resist bending to some extent, it is assumed that the resistance from a single mechanism is negligible; bending stiffness is added by using pairs of mechanisms, as circled on the segment in (a). In the model, each pair of mechanisms is represented by a set of four collocated single-degree-of-freedom damped springs (b), capturing the resistance to bending and translation. [Note that in (b), each set of four collocated springs is drawn as a single spring-damper pair].

mechanism is loaded and unloaded incrementally with known weights [Fig. 4(a)]. In the dynamic configuration, a cantilever suspends the magnet over the superconductor, with a parallel-blade flexure restricting the motion to a line parallel to the superconductor surface (Fig. 5). In both configurations, the superconductor rests on a Zerodur platform insulated by a Styrofoam box that contains the liquid nitrogen; areas outside this box are at room temperature except as cooled by stray nitrogen vapors or thermal conduction. For stability, the platform is supported by three Zerodur legs that pass through the Styrofoam to rest on a granite table. To allow for switching between configurations, the lever arm and cantilever are both removable, aside from their support posts, which are fixed to the platform. The lever arm, cantilever, and support posts are all constructed of Zerodur as well.

3.2 Static Measurements (Perpendicular Stiffness)

To determine the perpendicular stiffness and investigate the mechanism behavior, we measured changes in the magnet-superconductor separation as the flux-pinning mechanism was loaded with a sequence of known weights and incrementally unloaded. In these static experiments, the superconductor was constrained so that changes in the magnet-superconductor separation corresponded to displacements of the magnet. These displacements, Δh , were measured by reflecting a laser off a mirror mounted to a lever arm that rotated as the separation changed, and tracking the location of the reflected beam spot on a target a distance D away (Fig. 4). The raw measurements, then, consisted of a set of positions for the reflected beam spot. We determined these positions, H_{target} , by inking each location directly on the target and, after testing, using a drafting machine to meticulously measure the distance between each location and a reference line.

Processing this raw data requires relating changes in H_{target} to Δh , which can be accomplished by considering the experiment geometry and the changes that occur as the mechanism and apparatus transition from room temperature to cryogenic temperatures and as the spacer that constrains the magnet-superconductor separation during cooling is removed. At room temperature, the location of the reflected beam spot is

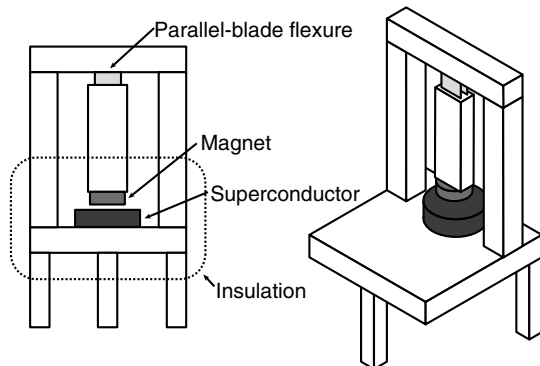


Fig. 5 The dynamic measurements (schematic). The parallel stiffness and damping are determined by measuring the impulse response. For these measurements, a cantilever suspends the magnet over the superconductor, with a parallel-blade flexure restricting the motion to a line parallel to the superconductor surface.

determined by the deviation of the mirror normal from 45 deg, α_n , the laser misalignment, α_l , and the initial tilt of the lever arm, α_p [Fig. 4(b)]. Since the pivot mirror rotates with the lever arm, the total tilt changes the height of the spot where the laser intersects the mirror, shifting the height of the reflected beam by δz . As liquid nitrogen is added, the apparatus expands/contracts on a global scale, changing the position of the reflected beam spot by Δz_{cool} . Similarly, contraction of the magnet, superconductor, and spacer alters the total height of the mechanism, causing the pivot arm to rotate by an additional amount θ_{cool} . These two effects are distinguished by comparing changes in the location of the reflected pivot beam spot to changes in the location of a reflected spot from a laser aimed at a reference mirror attached to a stationary part of the apparatus. After the experiment reaches thermal equilibrium, the spacer between the magnet and the superconductor is removed, and the mechanism is loaded with a known weight. As the magnet-superconductor separation decreases in response to the weight, the lever arm rotates by an additional amount θ_{weight} .

At each stage of the experiment, the translation of the reflected pivot beam spot is related to the total lever arm rotation, θ , by

$$D \tan(2\theta + 2\alpha_n - \alpha_l) = H_{\text{target}} - \Delta z_{\text{cool}} - \delta z, \quad (1)$$

where H_{target} is the height of the reflected pivot beam spot relative to the height of the point where the beam intersects the mirror when the experiment is at room temperature.[¶] Since δz is given by

$$\delta z = \left[\frac{\sin \theta(x_2 - x_1) + \cos \theta(z_2 - z_1)}{\cos \theta(x_2 - x_1) - \sin \theta(z_2 - z_1)} \right] (x_\alpha - x_1 \cos \theta + z_1 \sin \theta) + x_1 \sin \theta + z_1 \cos \theta - z_\alpha,$$

where (x_1, z_1) and (x_2, z_2) are points on the mirror surface and (x_α, z_α) is the point where the beam hits the mirror initially,[¶] Eq. (1) can be rewritten as

$$\begin{aligned} & D \tan(2\theta + 2\alpha_n - \alpha_l) \\ &= H_{\text{target}} - \Delta z_{\text{cool}} - x_1 \sin \theta - z_1 \cos \theta + z_\alpha \\ &\quad - \left[\frac{\sin \theta(x_2 - x_1) + \cos \theta(z_2 - z_1)}{\cos \theta(x_2 - x_1) - \sin \theta(z_2 - z_1)} \right] (x_\alpha - x_1 \cos \theta + z_1 \sin \theta). \end{aligned} \quad (2)$$

The various rotation angles are determined by iteratively solving Eq. (2) for θ and noting that

[¶]This equation can be derived by using Snell's law and considering the experiment geometry [Fig. 4(b)]. The angle between the reflected beam and horizontal, θ_w , is related to H_{target} and D by $D \tan \theta_w = H_{\text{target}} - \Delta z_{\text{cool}} - \delta z$. This angle also simultaneously satisfies the equations $(\pi/2) = \theta_w + 2\theta_l - \alpha_l$ and $(\pi/4) = \theta_w + \theta_l - \theta - \alpha_n$, where θ_l is the angle of incidence of the laser beam. Combining these three equations results in Eq. (1).

[¶]This expression for δz can be derived by considering how the coordinates of two points on the mirror change as the mirror rotates by θ , using these coordinates to define a line along the mirror surface, and finding the point on this line whose x -coordinate equals x_α .

$$\theta = \begin{cases} \alpha_p & \text{when the experiment is warm,} \\ \alpha_p + \theta_{\text{cool}} & \text{when the experiment reaches thermal equilibrium,} \\ \alpha_p + \theta_{\text{cool}} + \theta_{\text{weight}} & \text{when weights are applied.} \end{cases}$$

To relate θ_{weight} to the displacement caused by adding the weight, Δh_{weight} , we begin by noting that once the experiment reaches thermal equilibrium, the height of the lever arm pivot point relative to the bottom of the superconductor, H_{cold} , is a constant. Before the spacer is removed, H_{cold} is given by

$$d \sin(\alpha_p + \theta_{\text{cool}}) + l \cos(\alpha_p + \theta_{\text{cool}}) + h_{\text{disk}} - \Delta h_{\text{disk}} + h - \Delta h_{\text{cool}},$$

where d is the distance from the pivot point to the end of the lever arm, l is the length of the vertical section of the lever arm, h_{disk} is the thickness of the Zerodur disk at room temperature, Δh_{disk} is the change in h_{disk} due to cooling, h is the distance from the bottom of the superconductor to the top of the magnet at room temperature, and Δh_{cool} is the change in h due to cooling (Fig. 6). After the spacer is removed and weights are applied, H_{cold} is given by

$$d \sin(\alpha_p + \theta_{\text{cool}} + \theta_{\text{weight}}) + l \cos(\alpha_p + \theta_{\text{cool}} + \theta_{\text{weight}}) + h_{\text{disk}} - \Delta h_{\text{disk}} + h - \Delta h_{\text{cool}} - \Delta h_{\text{weight}}.$$

Equating these two expressions and solving for Δh_{weight} , we find that

$$\begin{aligned} \Delta h_{\text{weight}} &= d \sin(\alpha_p + \theta_{\text{cool}} + \theta_{\text{weight}}) - d \sin(\alpha_p + \theta_{\text{cool}}) \\ &\quad + l \cos(\alpha_p + \theta_{\text{cool}} + \theta_{\text{weight}}) \\ &\quad - l \cos(\alpha_p + \theta_{\text{cool}}), \end{aligned}$$

which for small angles reduces to

$$\Delta h_{\text{weight}} = d\theta_{\text{weight}}. \quad (3)$$

For our experiments, the target was placed a distance $D = 15,300$ mm away, while the distance from the pivot point to the end of the lever arm was $d = 109$ mm. As a result, changes in the magnet-superconductor separation were magnified by approximately a factor of 280; a 0.5-mm change in the separation, for example, caused the reflected beam spot to move 140 mm on the target. All of the misalignment angles and θ_{cool} were less than 1 deg ($\alpha_l < 0.8$ deg, $\alpha_n = 0.6$ deg, $|\alpha_p| < 1$ deg, and $\theta_{\text{cool}} < 0.5$ deg), and since the magnet-superconductor separation was at most 2 mm, θ_{weight} also did not exceed 1 deg. Consequently, the total rotation angle θ did not exceed 2.5 deg, justifying the small angle assumption.

To investigate the mechanism behavior, the mechanism was loaded with a sequence of weights, gradually increasing to a maximum of W_{seq} , then unloaded in reverse order. After several repetitions, W_{seq} was incremented to a new value. As Fig. 7 shows, the resulting displacement pattern is affected by the heaviest weight that has ever been applied, W_{max} . As W_{max} increases, the pattern shifts toward larger displacements. (If W_{seq} is less than W_{max} , the pattern does not shift back toward smaller displacements.) This effect appears to be an offset only, with no effect on the stiffness; when the offset between the average initial displacements for two sequences is subtracted, the data points for both sequences fall on the same curve. For practical purposes, this hysteretic behavior has two implications: the mechanisms can be trained to have a particular displacement pattern if the maximum load remains below some threshold,

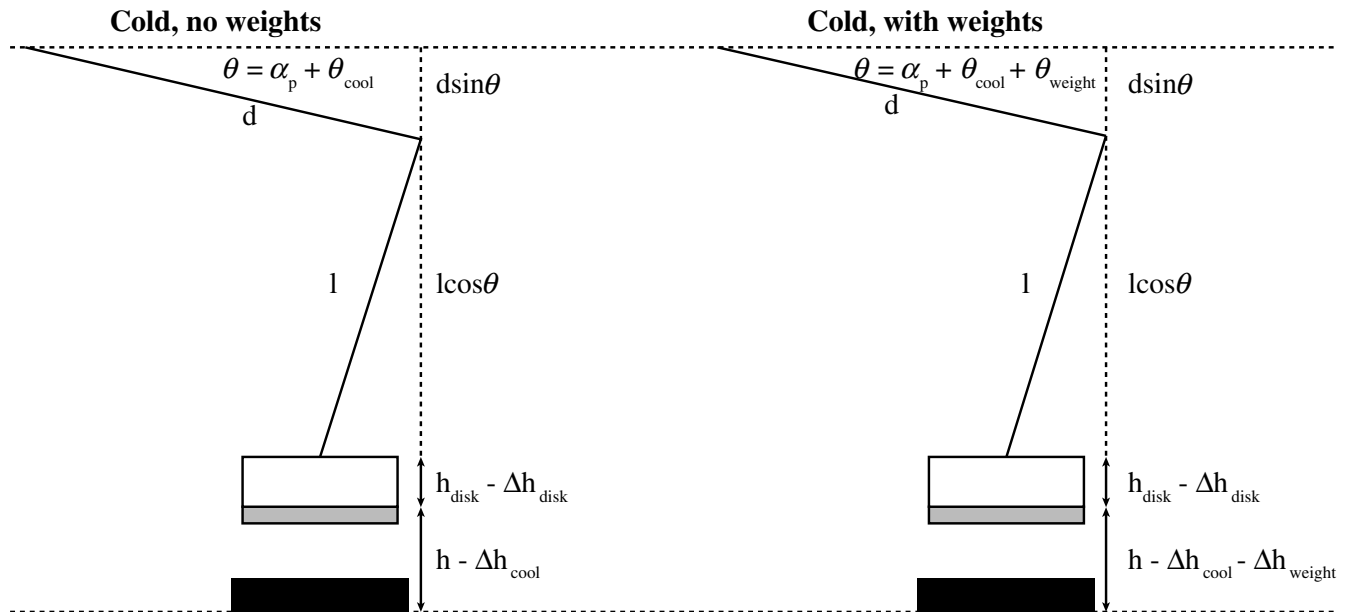


Fig. 6 Relating θ_{weight} to the magnet displacement. Since the height of the lever arm pivot point relative to the bottom of the superconductor is a constant, the amount of rotation caused by adding weight, θ_{weight} , and the corresponding magnet displacement, Δh_{weight} , are related by $\Delta h_{\text{weight}} = d\theta_{\text{weight}}$ for small angles.

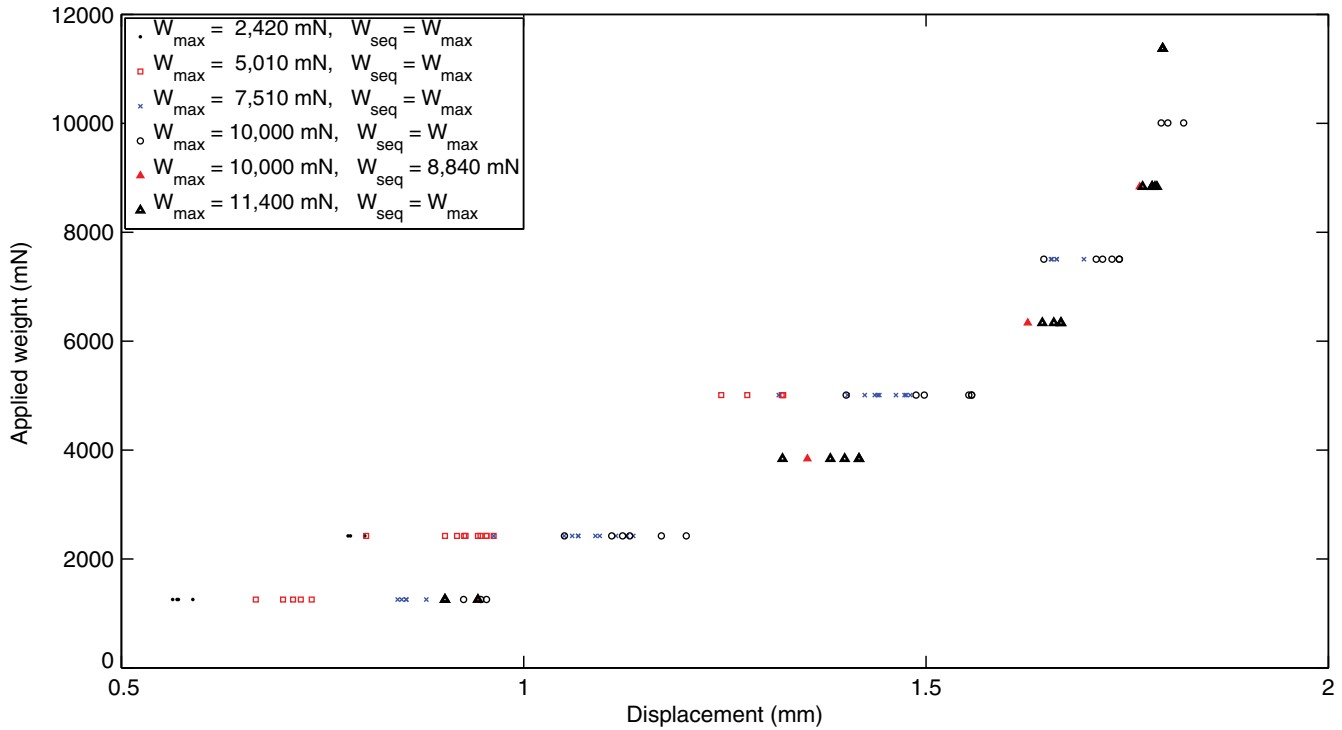


Fig. 7 The importance of W_{\max} . As the mechanism is sequentially loaded or unloaded, the displacement pattern depends on the maximum weight that has ever been applied, W_{\max} , rather than the heaviest weight in the sequence, W_{seq} . Increasing W_{\max} shifts the pattern toward larger displacements without affecting the stiffness, and the pattern does not shift back toward smaller displacements if $W_{\text{seq}} < W_{\max}$. As an example, the displacement data from a sample experiment is shown, with different symbols representing various weight sequences. For the initial sequence, one weight is applied and removed repeatedly, causing the displacement to oscillate between two values. For the second sequence, two weights are applied in order, then removed. Initially, the displacement pattern is the same as for the first sequence, but once the applied weight exceeds the previous value of W_{\max} , the displacements change, following a different pattern as the weights decrease. This new pattern is then followed (whether weights are added or removed) until a subsequent sequence increases W_{\max} .

and the displacements will increase if this threshold is exceeded. Since the mechanisms are situated along the segment edges, the ability to specify the displacement pattern corresponds to an ability to specify the size of the gap between adjacent segments, which may prove beneficial in situations where the gap must remain above a minimum value.

The stiffness values can be estimated from the displacement pattern by approximating the derivative between adjacent data points. For two points $(\Delta h_{\text{weight},1}, W_1)$ and $(\Delta h_{\text{weight},2}, W_2)$, the stiffness corresponding to the mean displacement is approximately

$$k = \frac{W_2 - W_1}{\Delta h_{\text{weight},2} - \Delta h_{\text{weight},1}}.$$

As Fig. 8 shows, the stiffness is nonlinear, and the shape of the curve depends on both the cooling separation and the magnet displacement. Decreasing the cooling separation increases the stiffness, as expected, and also leads to smaller shifts as W_{\max} increases. For a fixed cooling separation, the stiffness varies by an order of magnitude as the magnet-superconductor separation diminishes, increasingly sharply as the separation approaches zero. This increase is to be expected; previous work suggests that the stiffness increases nearly exponentially as the magnet approaches the superconductor.^{30,31} This behavior may prove beneficial in an edgewise-connected mirror: the

mechanisms provide increasing resistance the closer the two segments come to colliding.

3.3 Dynamic Measurements (Parallel Stiffness and Damping)

To determine the parallel stiffness and damping and study their dependence on the mechanism implementation, we measured the impulse response for mechanisms with various magnets and cooling separations of 0.5–2 mm. Since flux pinning is affected by the shape and strength of the magnetic field present during cooling, magnets of different diameter and thickness were tested. In addition to conventional magnets, we also tested a Swirl magnet, a neodymium magnet imprinted with a magnetic field pattern that preferentially allows rotational motion.^{32–34} This pattern was selected in order to investigate whether the rapidly changing magnetic field and its resistance to translation correspond to an increase in k_{\parallel} . For each mechanism, the stiffness and damping were extracted from the impulse response using eigensystem realization analysis,³⁵ and the results of multiple trials were averaged.

As shown in Figs. 9 and 10, k_{\parallel} and c generally increase as the cooling separation decreases or the magnet strength increases. These trends are to be expected since both stronger magnets and smaller cooling separations increase the magnetic flux penetrating into the superconductor during cooling; previous

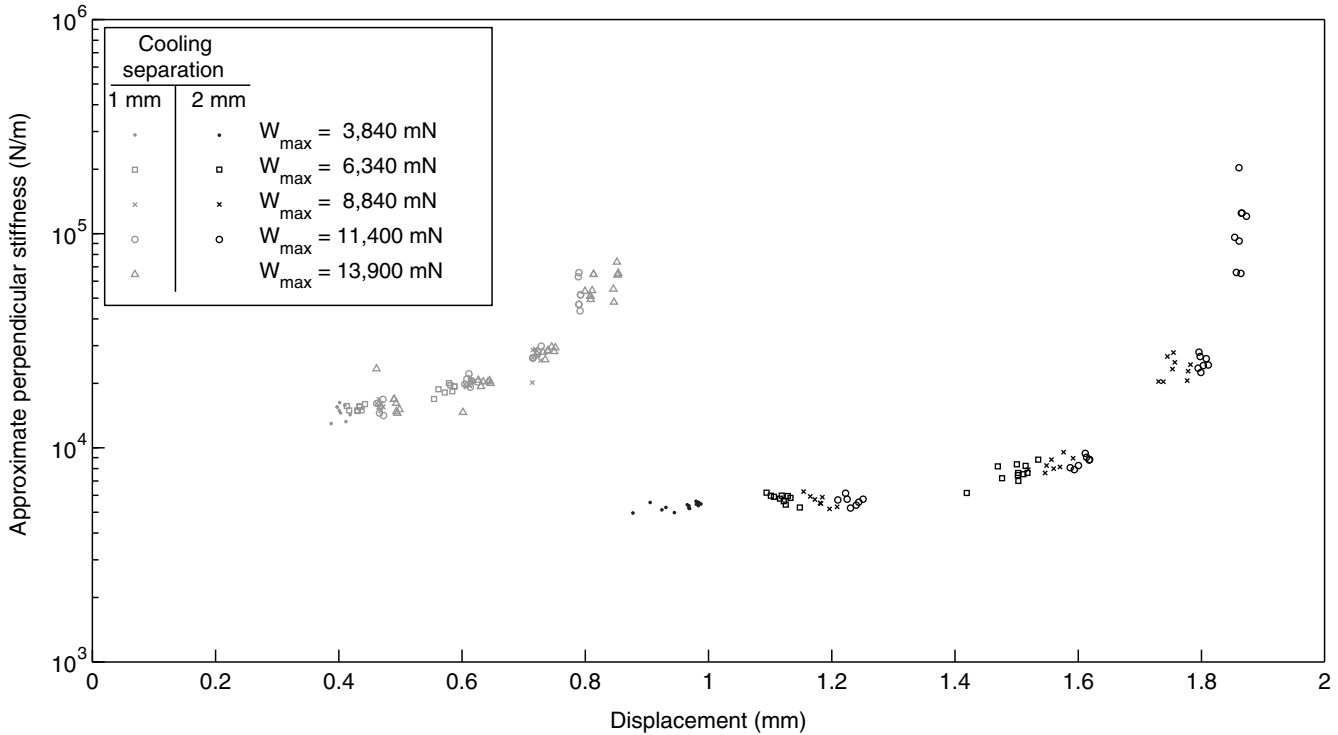


Fig. 8 The perpendicular stiffness. As the magnet approaches the superconductor, the stiffness varies by an order of magnitude, increasing sharply as the magnet-superconductor separation vanishes. As a result, the flux-pinning mechanism provides increasing resistance the closer two segments in an edge-wise-connected mirror come to colliding. Decreasing the cooling separation leads to stiffer mechanisms.

measurements have suggested that the stiffness increases nearly exponentially as the cooling separation decreases.^{16,36} For the mechanisms tested, the effects of varying the cooling separation are particularly noteworthy. As the cooling separation decreased from 2 to 0.5 mm, the stiffness increased by a factor of 2–10, with typical values on the order of 1000 N/m, and the damping increased by up to an order of magnitude, with typical values on

the order of 1–10 kg/s. By comparison, previous measurements using comparable mechanisms with a cooling separation of 5 cm reported stiffnesses on the order of 7 N/m and no discernible damping.¹⁶

While additional testing is needed to investigate the merits of using patterned rather than conventional magnets, the Swirl magnet considered in these experiments seems less useful

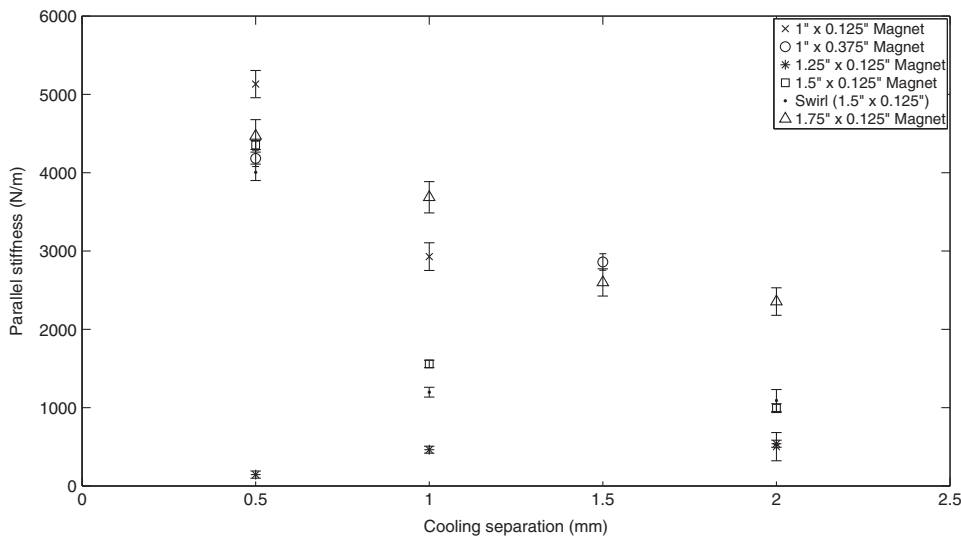


Fig. 9 The parallel stiffness for various mechanisms. In general, stronger magnets and smaller cooling separations lead to stiffer mechanisms. With cooling separations on the order of 1 mm, the mechanisms tested typically have stiffnesses on the order of 1000 N/m. [Note that in some cases, the error bars are smaller than the symbols.] By comparison, previous measurements for similar mechanisms with cooling separations on the order of 5 cm reported stiffnesses on the order of 7 N/m.¹⁶

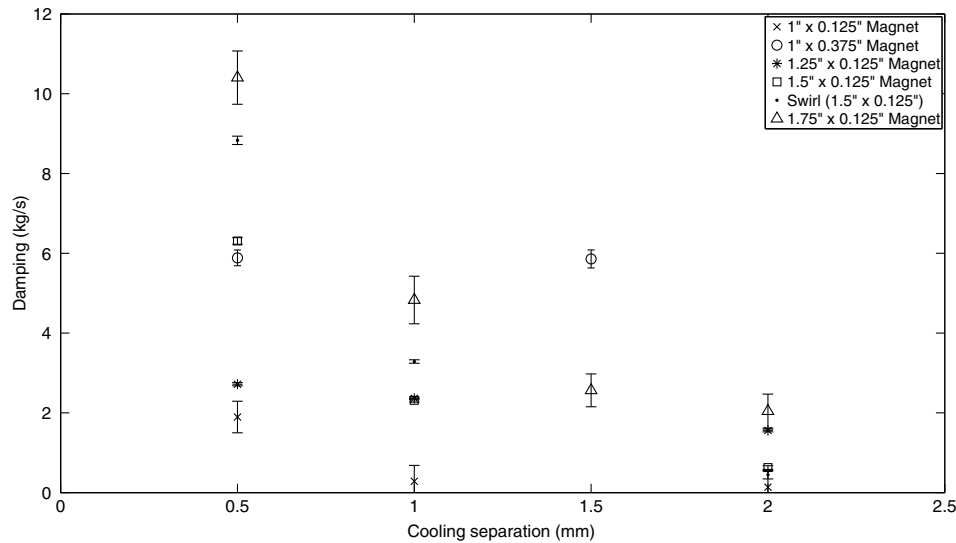


Fig. 10 The damping for various mechanisms. With cooling separations on the order of 1 mm, the mechanisms tested typically have damping values on the order of 1–10 kg/s. [Note that in some cases, the error bars are smaller than the symbols.] Smaller cooling separations and stronger magnets generally correspond to higher damping. By comparison, previous measurements with similar mechanisms and cooling separations on the order of 5 cm reported no discernible damping.¹⁶

than a conventional magnet of the same size. Although the Swirl pattern is designed to resist translation, the mechanism using the Swirl magnet typically had a lower stiffness than a mechanism using a conventional neodymium magnet with the same dimensions. Since the Swirl pattern concentrates the magnetic field in the near field,^{34,37} it is possible that less flux penetrates into the superconductor, leading to a lower stiffness. This effect could also influence the amount of damping. Although the mechanism using the Swirl magnet had higher damping than the mechanism using the conventional magnet, the reverse may be true if the mechanisms are modified to include nonmagnetic conductive materials: since the increased damping is due to eddy current damping, the lower flux penetration of the Swirl magnet could correspond to smaller gains.

Although investigating the relationship between mechanism properties, such as the cooling separation and choice of magnet, and the resulting stiffness and damping values provides insight into how to design mechanisms with the desired values, an important result of these measurements is determining the order of magnitude for the stiffness and damping values that can be achieved with inexpensive magnets and cooling separations on the order of 1 mm. Previous modeling work¹³ has shown that the mechanism stiffness determines whether the segments of an edgewise-connected mirror act as individual rigid bodies or as a cohesive unit. Changing the stiffness by orders of magnitude (by reducing the cooling separation from 5 cm to 1 mm, for example) can therefore alter the fundamental mirror behavior. We will discuss the behavior of an edgewise-connected mirror using flux-pinning mechanisms in Sec. 4.

4 Simulated Mirror Performance

To investigate the performance improvements provided by placing flux-pinning mechanisms along the segment edges, we consider the impulse response of a 15-m edgewise-connected mirror composed of two rings of hexagonal segments. The choice of a 15-m mirror is motivated by the preliminary results of the AURA “Beyond JWST” study, which indicate that while the

minimum acceptable aperture diameter for the next generation of space telescopes is 6.5 to 8 m, a 12- to 14-m aperture is desired, and a 16-m aperture is highly desirable.³⁸ Six pairs of flux-pinning mechanisms are placed along each segment edge, with stiffness and damping values representative of the measurements presented in Sec. 3. As a basis for comparison, we consider two additional mirrors: a monolithic mirror with the same size, shape, and material properties as the edgewise-connected mirror; and an edgewise-connected mirror that is identical to the one described above except that the mechanisms have no damping. All three mirrors are mounted identically, and for simplicity, they are kinematically mounted at three points. Consequently, the mechanisms serve as the only connections between the segments of an edgewise-connected mirror; the segments are not also connected via a backplane. Although it is likely that the segments of an edgewise-connected mirror would be mounted individually to a backplane in practice, the problem of segmented mount design is beyond the scope of this paper.

The impulse response for each mirror is determined using a parametric finite-element model. Previous work with this model¹³ has shown that the magnitude of the impulse response is affected primarily by the mechanism stiffness. The mechanism stiffness affects the strength of the connections between the segments, determining whether the segments of an edgewise-connected mirror behave as individual rigid bodies or as a single unit. When the mechanisms are sufficiently stiff that the total bending stiffness along a segment edge is comparable to the bending stiffness of the monolithic mirror, the edgewise-connected mirror behaves similarly to the monolith. In this case, the magnitude of the impulse response for the edgewise-connected mirror is comparable to that of the monolith since disturbances propagate similarly across either mirror. For lower-stiffness mechanisms, the connections between the segments can be much weaker, and the segments tilt as individual rigid bodies rather than bending together. In this case, the magnitude of the disturbance response can be much lower since disturbances

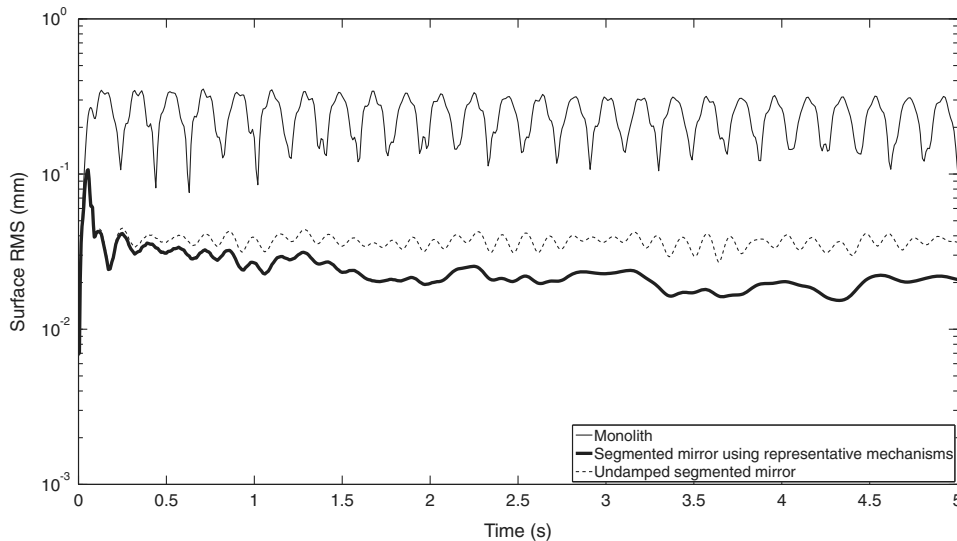


Fig. 11 The impulse response for an edgewise-connected mirror with flux-pinning mechanisms. The damping provided by the flux-pinning mechanisms improves the disturbance response, reducing the number of vibrating frequencies and increasing the decay rate. For reference, the responses of the undamped segmented mirror and a monolith of the same size and shape are also shown; the segmented mirrors are identical except for the mechanism damping. The responses of the segmented mirrors are approximately an order of magnitude lower than the response of the monolith since disturbances do not propagate as effectively across a primary consisting of weakly connected segments; the effects of a backplane are not considered in this model.

do not propagate as effectively across the edgewise-connected mirror. (As a reminder, the model does not consider the effects of a backplane).

For the example 15-m edgewise-connected mirror, the mechanism stiffness must be on the order of 10^7 N/m in order for the mirror to behave similarly to the monolith.¹³ Since this stiffness is approximately four orders of magnitude larger than the stiffnesses measured in Sec. 3, the mechanisms would need to be stiffened substantially, which may be achievable by using smaller cooling separations or stronger magnets. Although the tested mechanisms are not sufficiently stiff to structurally connect the segments of the example edgewise-connected mirror, their stiffness contributions may lessen the requirements for the mirror support structure.

A particularly interesting application for the tested mechanisms is providing damping to a cryogenic mirror. As shown in Fig. 11, the damping contributions from the mechanisms improve the impulse response of the edgewise-connected mirror, reducing the number of oscillating frequencies and increasing the decay rate. Though the improvements are modest, more substantial gains are possible with higher damping values.¹³ These values may be attainable by placing nonmagnetic conductive materials near the mechanisms. Approaches include placing bulk material adjacent to the mechanisms and fabricating the mirror segments from a nonmagnetic conductive material. The amount of additional damping will likely depend on a variety of factors including the magnet strength, the distance between the material and the moving magnet, the material conductivity, and the amount of material. Quantifying how the amount of material and its placement affect the damping is a subject for future investigations.

5 Summary

As future astrophysics missions require larger far-IR to sub-millimeter space telescopes, maintaining the stability of the

cryogenic primary becomes increasingly challenging. One approach to increasing the mirror stiffness and damping is to use an edgewise-connected architecture, with flux-pinning mechanisms placed along the segment edges. Consisting of a configuration of magnets and superconductors, flux-pinning mechanisms are uniquely suited for cryogenic mirrors since they require low temperatures to operate, unlike mechanical devices, which can have problems with lubrication, CTE matching, and thermal snap. Like flexures, flux-pinning mechanisms preferentially allow motion in specific degrees of freedom, which depend on the mechanism design. These noncontacting mechanisms are passively stable and require no power other than the amount needed for cooling.

The stiffness and damping contributions from the flux-pinning mechanisms improve the mirror stability and lessen the requirements for the mirror support structure. As an example, we considered a flux-pinning mechanism consisting of a single magnet and superconductor. To examine how this type of mechanism can improve the performance of a sample 15-m mirror, we first measured the mechanism stiffness and damping using a specialized apparatus that we constructed out of Zerodur in order to solve the challenges presented by the need for cryogenic temperatures and nonmagnetic, nonconductive materials, which would not interact with the mechanism. We then entered the measurements into a parametric finite-element model to determine the resulting mirror behavior.

With an inexpensive magnet and a cooling separation on the order of 1 mm, our mechanisms have typical stiffness and damping values on the order of 5000 N/m and 5 kg/s, respectively. With these values, the mechanisms provide modest improvements to the mirror performance, increasing the stiffness and decreasing the settling time. Greater stiffnesses can be achieved by using stronger magnets or smaller cooling separations, and the damping can be adjusted independently by placing nonmagnetic conductive materials near the mechanism. Quantifying the

increases in damping due to the material amount and placement remains a subject for future investigation.

Acknowledgments

The authors would like to thank the NASA Graduate Student Researchers Program for its support of this project, and the following people whose invaluable assistance made this project possible: Charlie Griffith, for teaching the fine art of grinding and polishing Zerodur; Bob Engberg and his team, for their support of the vibration measurements; and Ephraim Garcia, for his insights into processing the vibration data. This project was funded by NASA Grant Nos. NNX09AJ18H and NNX12AC61G.

References

- D. Leisawitz, "NASA's far-IR/submillimeter roadmap missions: SAFIR and SPECS," *Adv. Space Res.* **34**(3), 631–636 (2004).
- H. A. Thronson et al., "Astronomy enabled by Ares V heavy lift," *Future In-Space Operations White Paper* (2007).
- D. J. Benford et al., "Mission concept for the Single Aperture Far-Infrared (SAFIR) Observatory," *Astrophys. Space Sci.* **294**, 177–212 (2004).
- G. H. Rieke et al., "Charting the winds that change the universe, II. The Single Aperture Far Infrared Observatory (SAFIR)," in *Proc. Second Workshop on New Concepts for Far-Infrared and Submillimeter Space Astronomy*, D. J. Benford and D. T. Leisawitz, Eds., pp. 157–166, NASA, Greenbelt, MD (2004).
- J. C. Mather, "Prospects for future observations in the mid/far IR," *AIP Conf. Proc.* **666**, 347–354 (2003).
- G. L. Pilbratt, "Herschel mission overview and key programmes," *Proc. SPIE* **7010**, 701002 (2008).
- L. Feinberg et al., "Space telescope design considerations," *Opt. Eng.* **51**(1), 011006 (2012).
- A. Danjon and A. Couder, *Lunettes et télescopes: Théorie, conditions d'emploi, description, réglage, histoire*, p. 570, Editions de la Revue d'Optique Théorique et Instrumentale, Paris (1935).
- M. Levine and C. White, "Material damping experiments at cryogenic temperatures," *Proc. SPIE* **5179**, 165–176 (2003).
- O. Romberg et al., "Passive damping of spacecraft sandwich panels," in *Proc. 10th European Conf. on Spacecraft Structures, Materials, & Mechanical Testing*, pp. 1–8 (2007).
- P. Y. Bely, Ed., *The Design and Construction of Large Optical Telescopes*, Springer-Verlag, New York (2003).
- H. P. Stahl and L. Feinberg, "Summary of NASA advanced telescope and observatory capability roadmap," in *2007 IEEE Aerospace Conf.*, pp. 1–11 (2007).
- J. Gersh-Range, W. R. Arnold, Sr., and H. P. Stahl, "Edgewise connectivity: an approach to improving segmented primary mirror performance," *J. Astron. Telesc. Instrum. Syst.* **1**(1), in press (2014).
- Can Superconductors, "Superconducting YBCO levitation disks," http://shop.can-superconductors.com/attachment.php?id_attachment=2 (29 March 2012).
- A. Kordyuk, "Magnetic levitation for hard superconductors," *J. Appl. Phys.* **83**(1), 610–612 (1998).
- J. P. Shoer and M. A. Peck, "Flux-pinned interfaces for the assembly, manipulation, and reconfiguration of modular space systems," *J. Astronaut. Sci.* **57**(3), 667–688 (2009).
- L. C. Davis, "Lateral restoring force on a magnet levitated above a superconductor," *J. Appl. Phys.* **67**(5), 2631–2636 (1990).
- R. Williams and J. R. Matey, "Equilibrium of a magnet floating above a superconducting disk," *Appl. Phys. Lett.* **52**(9), 751–753 (1988).
- J. R. Hull and A. Cansiz, "Vertical and lateral forces between a permanent magnet and a high-temperature superconductor," *J. Appl. Phys.* **86**(11), 6396–6404 (1999).
- S. A. Basinger, J. R. Hull, and T. M. Mulcahy, "Amplitude dependence of magnetic stiffness in bulk high-temperature superconductors," *Appl. Phys. Lett.* **57**(27), 2942–2944 (1990).
- E. H. Brandt, P. Esquinazi, and H. Neckel, "A superconducting vibrating reed applied to flux-line pinning. I. Theory," *J. Low Temp. Phys.* **63**(3–4), 187–214 (1986).
- R. Grosser et al., "Vortex motion in superconducting $\text{YBa}_2\text{Cu}_3\text{O}_{7-x}$ inferred from the damping of the oscillations of a levitating magnetic microsphere," ArXiv e-prints arXiv: cond-mat/9901085 (1999).
- J. Shoer et al., "Microgravity demonstrations of flux pinning for station-keeping and reconfiguration of CubeSat-sized spacecraft," *J. Spacecr. Rockets* **47**(6), 1066–1070 (2010).
- J. Shoer and M. Peck, "Reconfigurable spacecraft as kinematic mechanisms based on flux-pinning interactions," *J. Spacecr. Rockets* **46**(2), 466–469 (2009).
- H. P. Stahl, M. Postman, and W. S. Smith, "Engineering specification for large-aperture UVO space telescopes derived from science requirements," *Proc. SPIE* **8860**, 886006 (2013).
- R. Nalbandian and A. E. Hatheway, "Extra Large Telescope Actuator (ELTA)," *Proc. SPIE* **4837**, 814–820 (2003).
- R. M. Warden, "Cryogenic nano-actuator for JWST," in *Proc. 38th Aerospace Mechanisms Symposium*, pp. 239–252, NASA, Hampton, VA (2006).
- L. Zago, P. Schwab, and D. Gallieni, "Development and testing of a high-precision, high-stiffness linear actuator for the focus-center mechanism of the SOFIA secondary mirror," *Proc. SPIE* **4014**, 392–398 (2000).
- SCHOTT North America, "Zerodur: zero expansion glass ceramic," http://www.us.schott.com/advanced_optics/english/download/schott_zerodur_katalog_july_2011_us.pdf (10 October 2013).
- P. Z. Chang et al., "Levitation force and magnetic stiffness in bulk high-temperature superconductors," *J. Appl. Phys.* **67**(9), 4358–4360 (1990).
- B. R. Weinberger, L. Lynds, and J. R. Hull, "Magnetic bearings using high-temperature superconductors: some practical considerations," *Supercond. Sci. Technol.* **3**(7), 381–388 (1990).
- Correlated Magnetics, "SwirlCode™ correlated pair," <http://www.correlatedmagnetics.com/products/swirlcodetm-correlated-pair/> (14 September 2013).
- Correlated Magnetics, "Programmed behavior," <http://www.correlatedmagnetics.com/technology/maglatch/> (14 September 2013).
- Correlated Magnetics, "Correlated magnetics," (2009).
- J. N. Juang and R. S. Pappa, "An eigensystem realization algorithm for modal parameter identification and model reduction," *J. Guid. Control Dyn.* **8**(5), 620–627 (1985).
- J. Shoer and M. Peck, "Stiffness of a flux-pinned virtual structure for modular spacecraft," *J. Br. Interplanet. Soc.* **62**, 57–65 (2009).
- Correlated Magnetics, "Coded magnetic structures and the shortest path effect," (2009).
- H. P. Stahl, private communication (2014).

Jessica Gersh-Range recently completed her PhD degree in mechanical engineering at Cornell University. She was the recipient of a NASA Graduate Student Researchers Program fellowship at Marshall Space Flight Center, and she has also worked at the Space Telescope Science Institute as a graduate student. She received her BA degree in physics from Swarthmore College in 2006, with a minor in mathematics. Her research interests include space optical systems, which combines her physics and engineering backgrounds.

H. Philip Stahl, senior optical physicist at NASA MSFC, is leading an effort to mature technologies for a large aperture telescope to replace Hubble. Previous assignments include developing JWST mirror technology. He is a leading authority in optical metrology, optical engineering, and phase-measuring interferometry. He is a member of OSA, SPIE (fellow), and 2014 SPIE president. He earned his PhD degree in optical science at the University of Arizona, Optical Sciences Center in 1985.

Biographies for the other authors are not available.